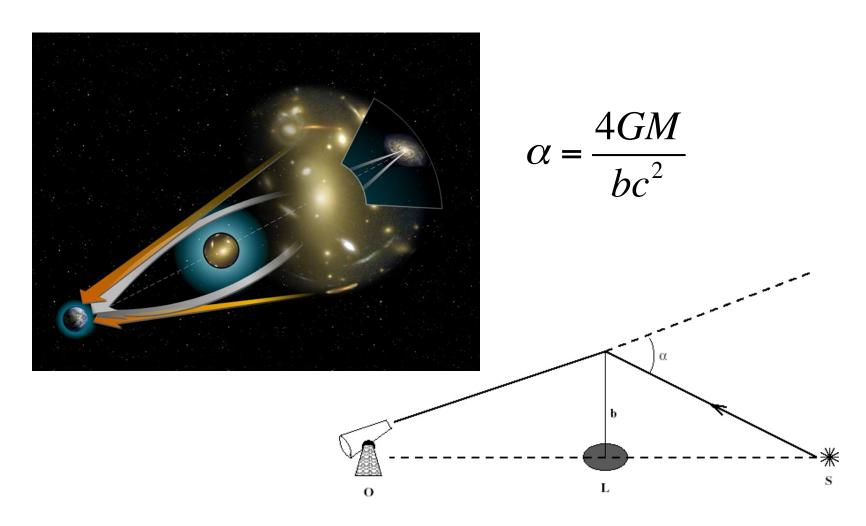
Constrain Point Spread function by using Galaxy Images

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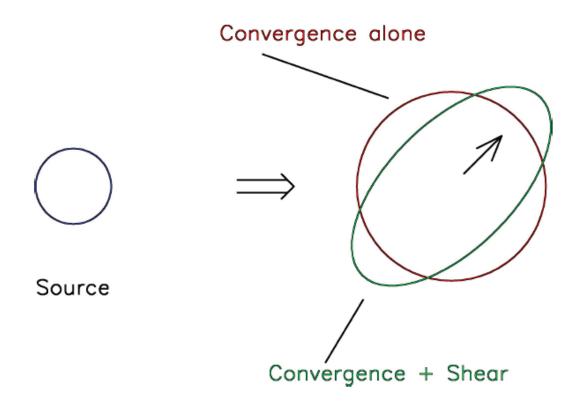
Outlines

- Quick view on cosmic shear measurement.
- Why we need galaxy images to constrain point spread function(PSF).
- How we do it.
- Some test results.

What is gravitational lensing?



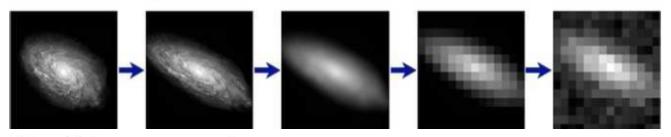
Weak lensing effects



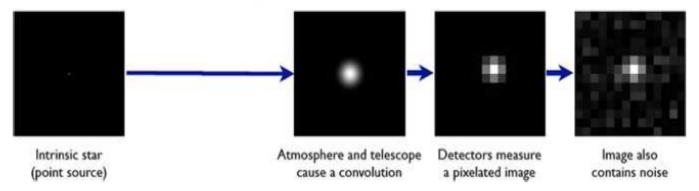
Point Spread Function

The Forward Process.

Galaxies: Intrinsic galaxy shapes to measured image:



Stars: Point sources to star images:



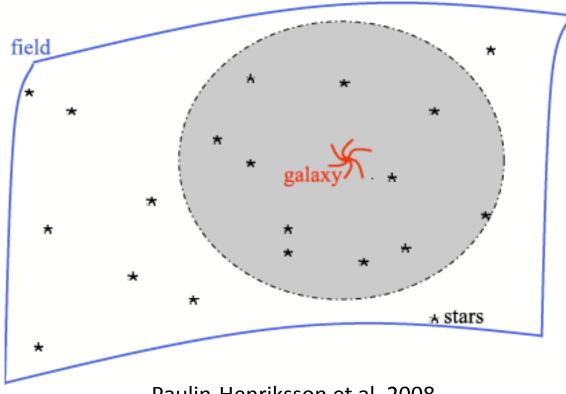
Usually, the distortion introduced by PSF is larger than that caused by cosmic shear!

How do we do the reconstruction?

- A) Star identification.
- B) Parameterized image of stars.
- C) Interpolate the parameters in the filed of view.
- D) Create the expected PSF at the position of galaxies.



Motivation



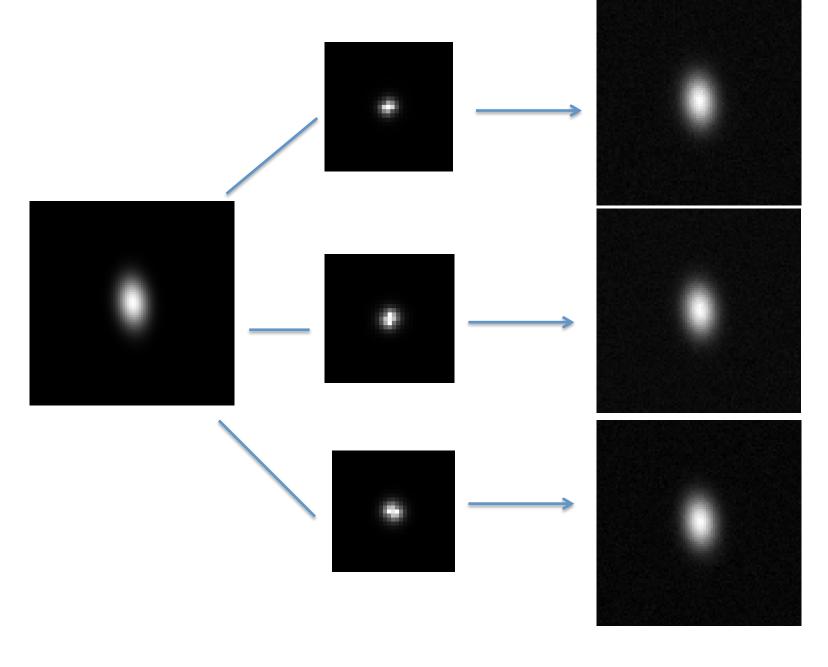
Paulin-Henriksson et al. 2008

$$\delta \epsilon^{\mathrm{sys}} \simeq \left(\epsilon_{\mathrm{gal}} - \epsilon_{\mathrm{PSF}} \right) \frac{\delta \left(R_{\mathrm{PSF}}^2 \right)}{R_{\mathrm{gal}}^2} - \left(\frac{R_{\mathrm{PSF}}}{R_{\mathrm{gal}}} \right)^2 \delta \epsilon_{\mathrm{PSF}}$$

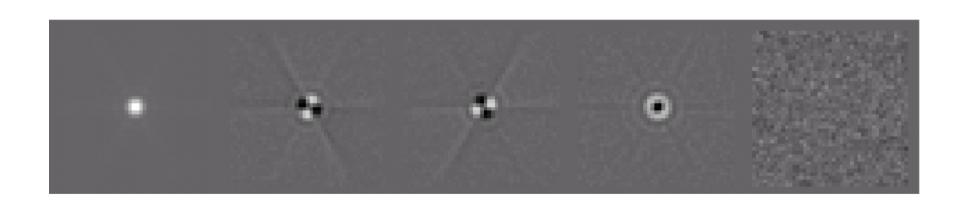
$$\sigma_{\rm sys}^2 \lesssim 10^{-7}$$



$$\frac{\sigma[R_{\rm PSF}^2]}{R_{\rm PSF}^2} \lesssim 10^{-3},$$
$$\sigma[\epsilon_{\rm PSF}] \lesssim 10^{-3}.$$



Unknows: NgxNg + nxNPgxNPg Knows: nxNgxNg

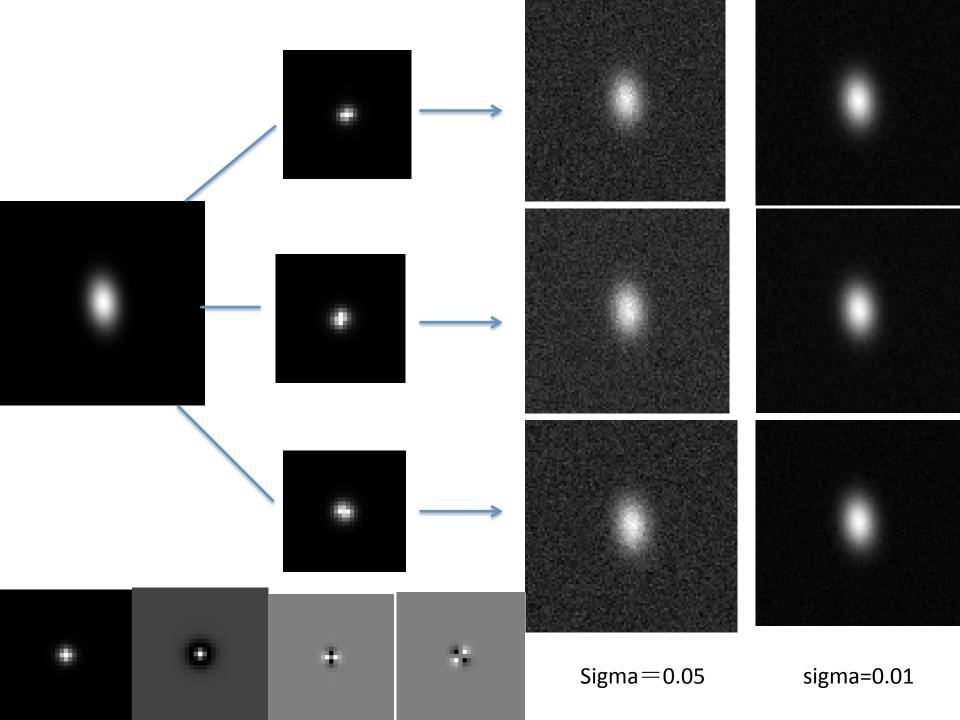


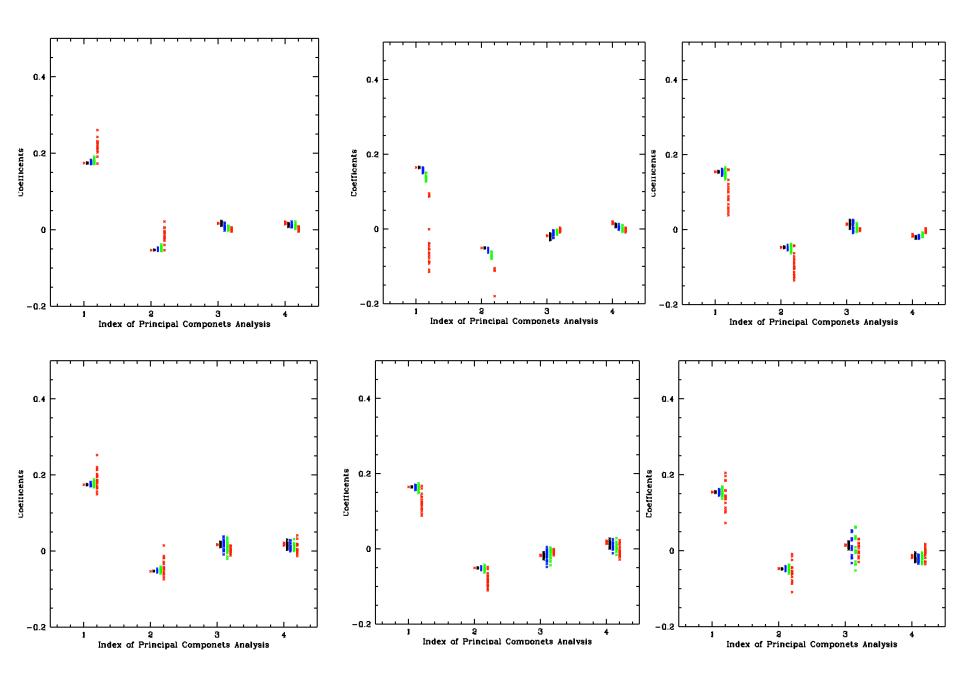
假设由其他星导出的PSF可以完备地描述星系所属的PSF

PSF的NgxNg个变量还可以用几个主成分的系数来表示, $PSF_i = \sum_{i} c_{il} PC_{il}$

$$G_0 \otimes PSF_i \otimes PSF_j = G_i \otimes PSF_j = G_j \otimes PSF_i = G_0 \otimes PSF_j \otimes PSF_i$$

$$\chi_{ij}^{2} = \sum_{k}^{N_{pixel}} \frac{(G_{i} \otimes PSF_{j} - G_{j} \otimes PSF_{i})_{k}^{2}}{\sigma_{ijk}^{2}}$$





Summary

- We constructed an algorithm to constrain PSF by using the galaxy images.
- Noise can bias the results but can be corrected with a simple filter.
- Apply to the real data soon.

Thank you!