

The segregation of baryons and dark matter during halo assembly

Liang Gao (NAOC)

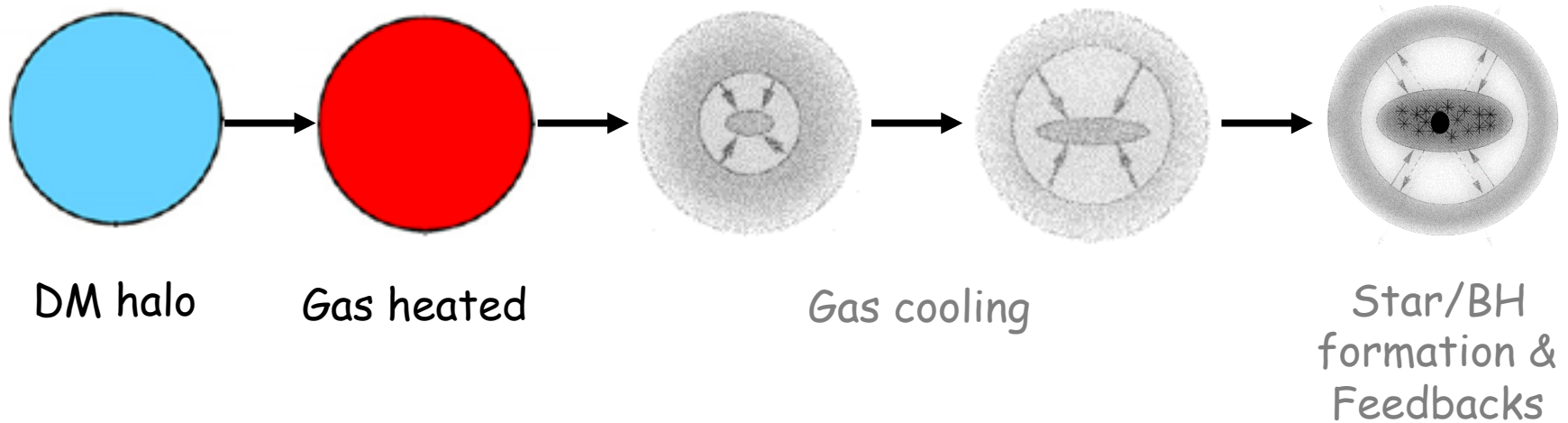
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Galaxy formation--well-mixing assumption

- Baryons and dark matter are well mixed before they are **segregated** due to *radiative cooling*.



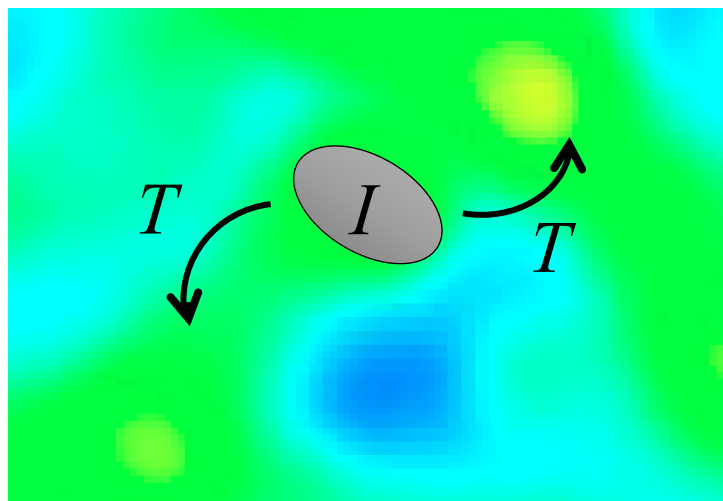
(Baugh 2006)

One important application: disk formation

The same spatial distribution (well-mixing) → the same tidal torques (Fall & Efstathiou 1980; Mo et al. 1998)

→ **identical specific angular momenta** ($\mathbf{j} = \mathbf{J}/M$)

→ calculating disk radii



$$\mathbf{j}_{\text{gas}} = \mathbf{j}_{\text{dm}}$$

Tidal torque theory

(Peebles 1969; Doroshkevich 1970; White 1984)

How good is this assumption?

Dark matter

$$\partial \mathbf{v} / \partial t + (\mathbf{v} \cdot \nabla) \mathbf{v} = -\nabla \phi$$

- Collisionless
- Gravity only
- Violent relaxation
- ...

Baryons (Gas)

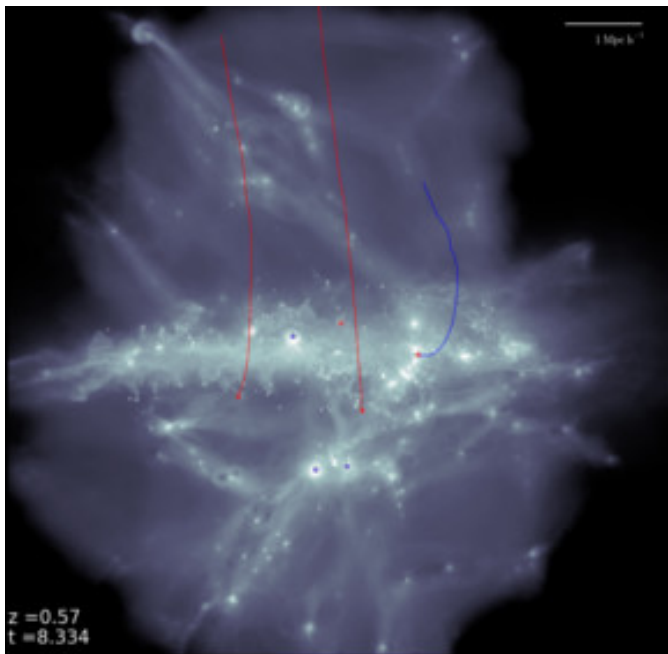
$$\partial \mathbf{v} / \partial t + (\mathbf{v} \cdot \nabla) \mathbf{v} = -1/\rho \nabla P - \nabla \phi$$

- Collisional
- Gravity + **pressure**
- Shocks
- ...

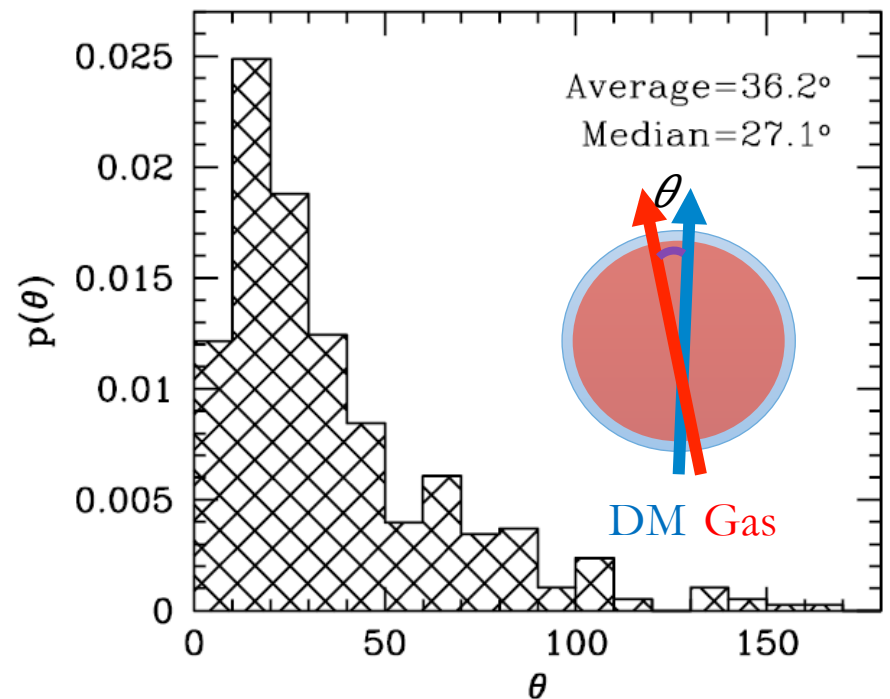
- It is not obvious that the well-mixing assumption should hold to high precision.

Evidences against the well mixing assumption

- “Cosmic web stripping” (Benitez-Llambay et al. 2013, 2016)
- Angular momentum misalignment (van den Bosch et al. 2002, 2003; Yoshida et al. 2003; Chen et al. 2003; Sharma & Steinmetz 2005; Croft et al. 2009; Bett et al. 2010; Zjupa & Springel 2016)



(Benitez-Llambay et al. 2013)



(van den Bosch et al. 2002)

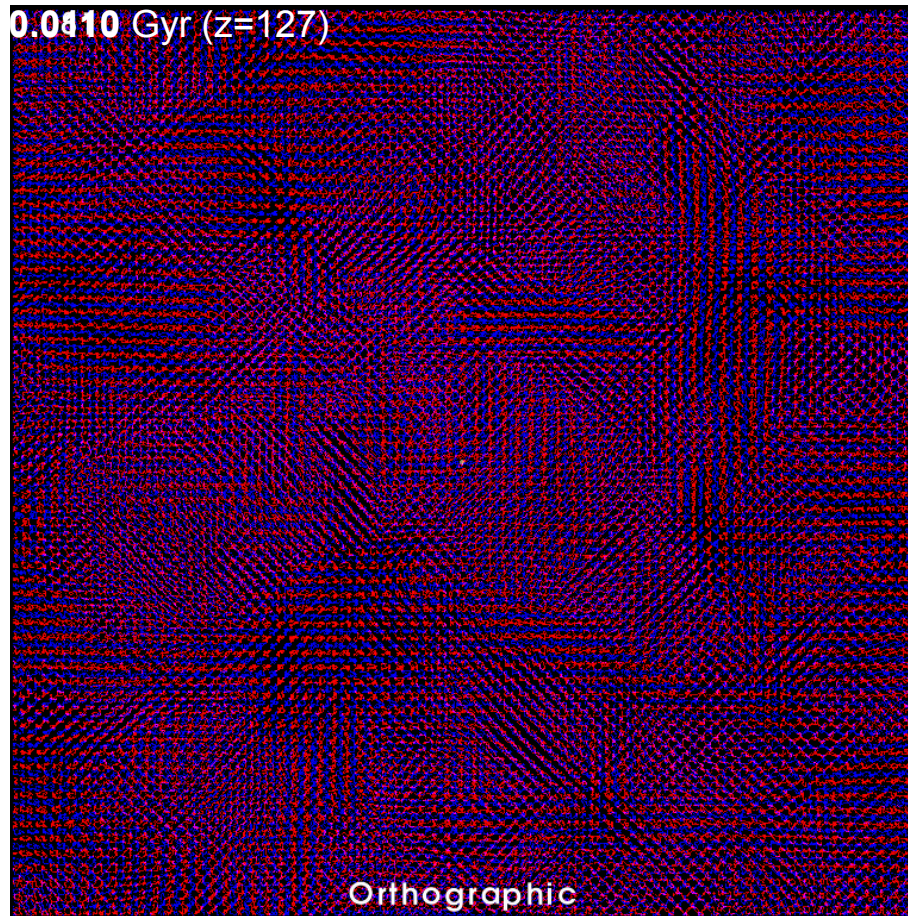
Questions

- Are DM and GAS component of a halo initially well-mixed or segregated?
- If segregated, any implications on current understanding of galaxy formation models? Any hints at the puzzling AMM between DM and gas?

Simulating halo assembly

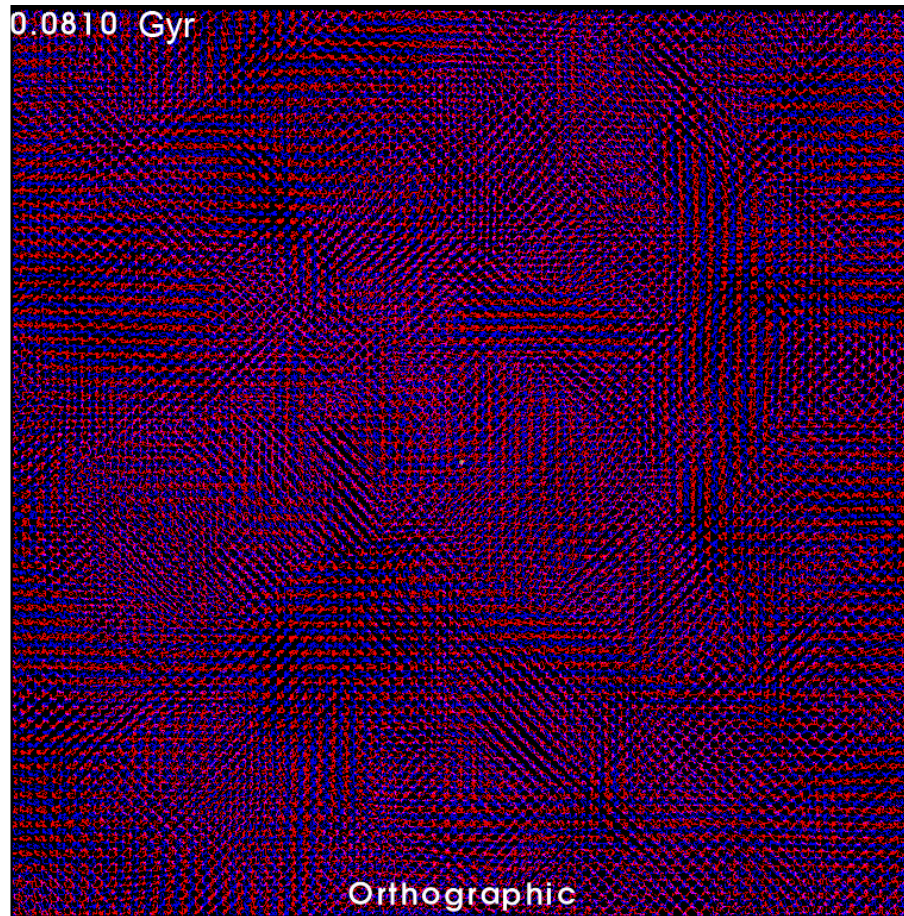
Blue: dark matter

Red: gas



Simulating halo assembly

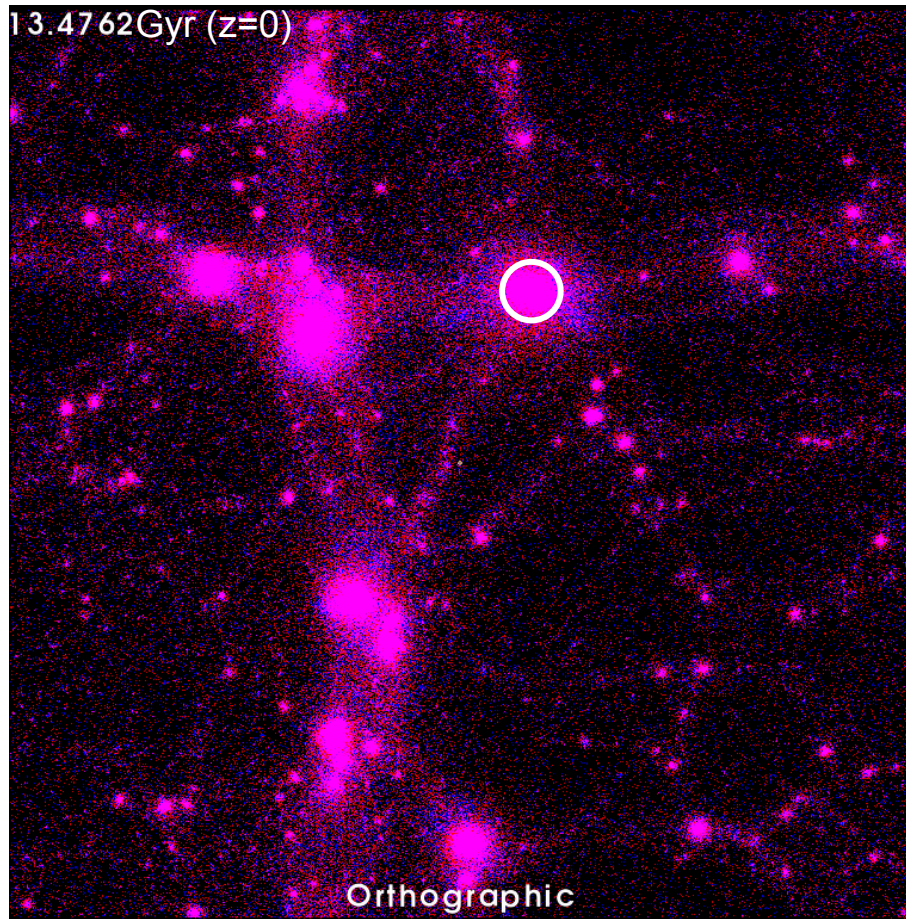
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Simulating halo assembly

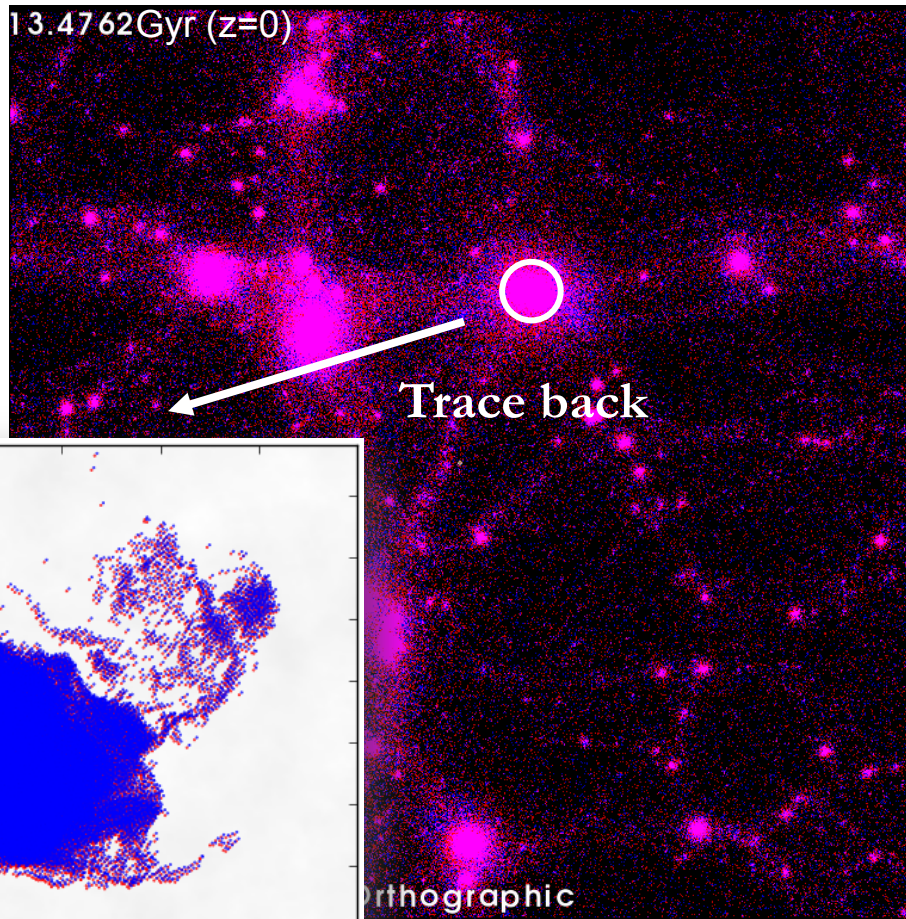
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Simulating halo assembly

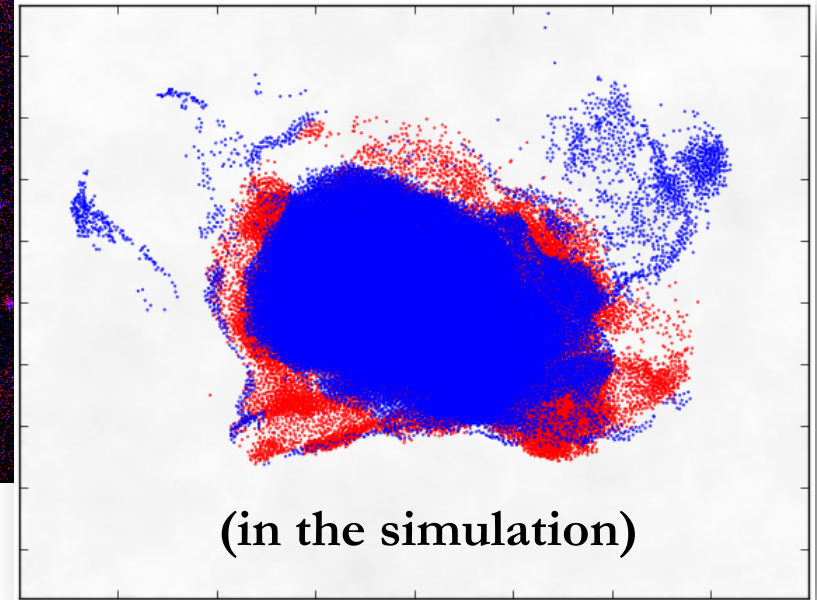
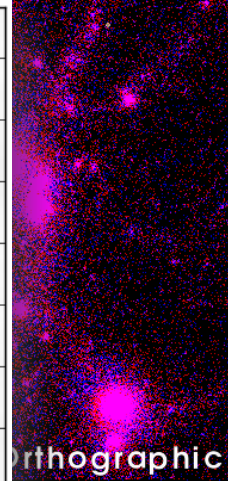
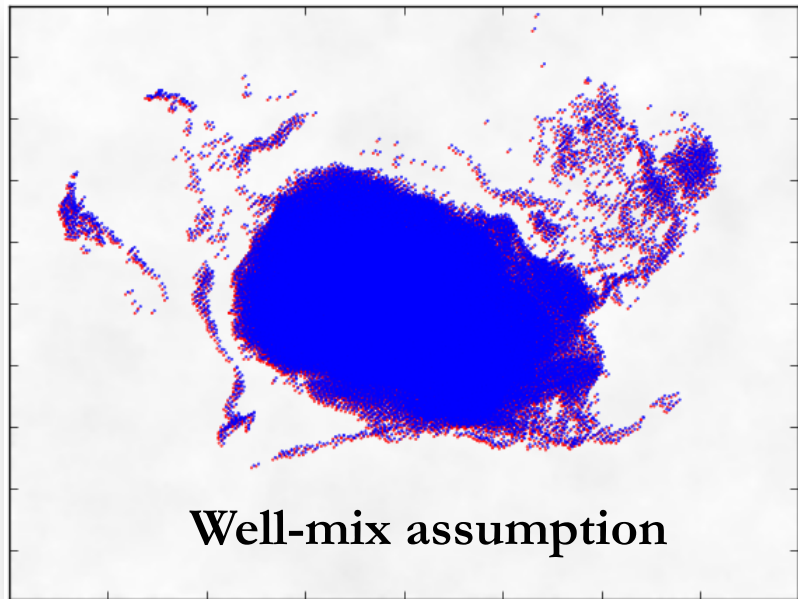
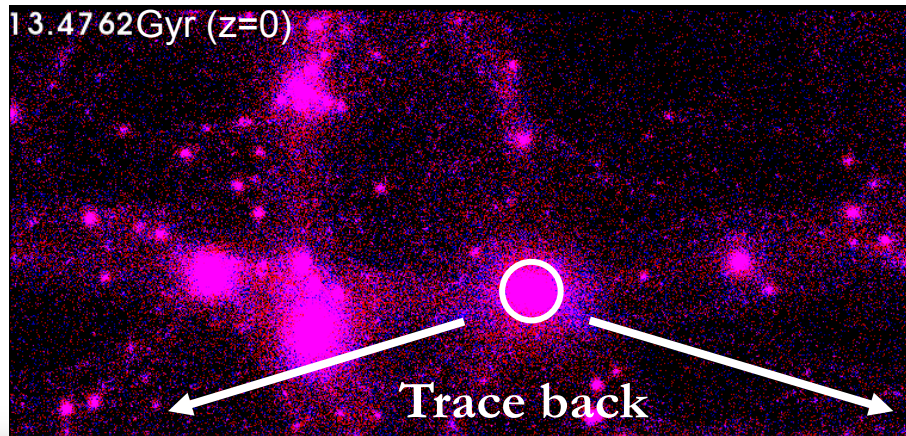
Blue: dark matter
Red: gas



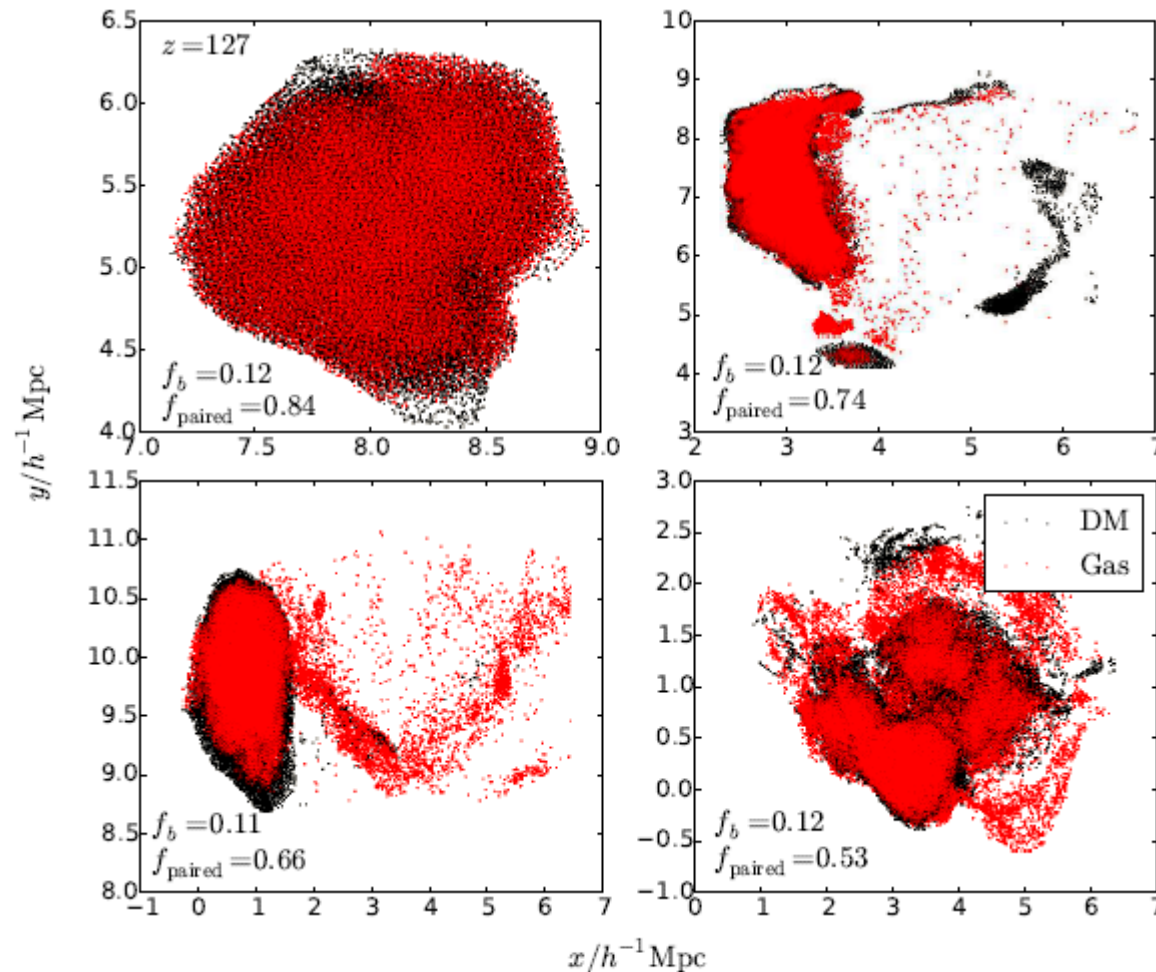
(if no segregation)

Simulating halo assembly

Blue: dark matter
Red: gas



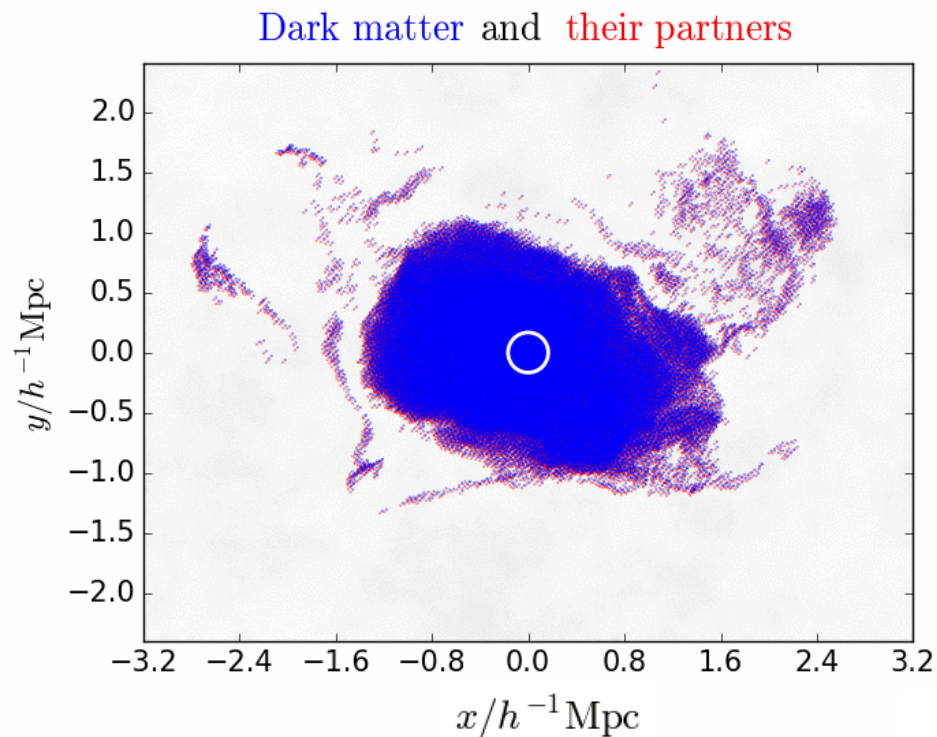
DM and gas of the present day haloes often do not come from the same regions in the initial conditions.



What causes DM-gas segregation?

- Merger
- Cosmic web striping
- Environmental effects
-

Example: halo merging can lead to segregation.

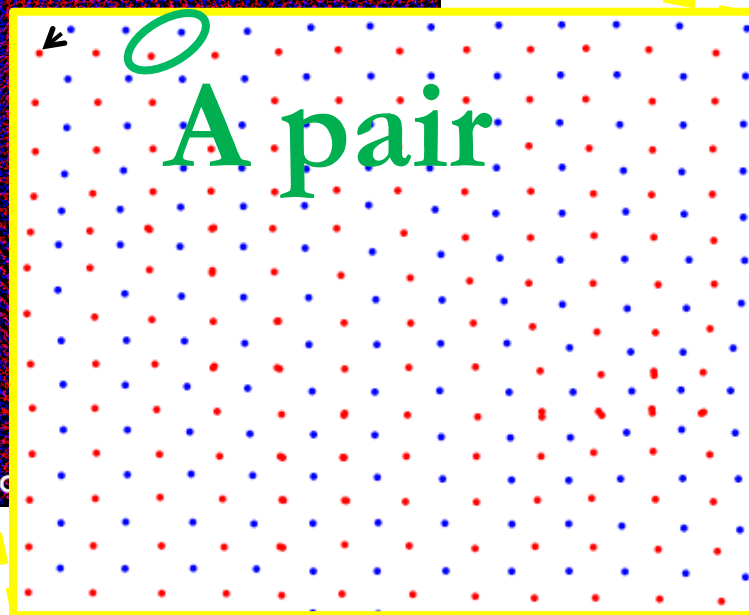
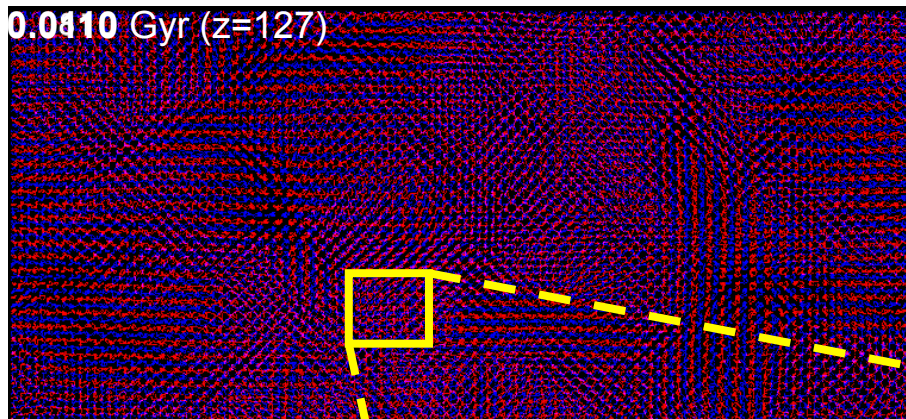


Simulating halo assembly

0.0410 Gyr (z=127)

Blue: dark matter

Red: gas



A pair

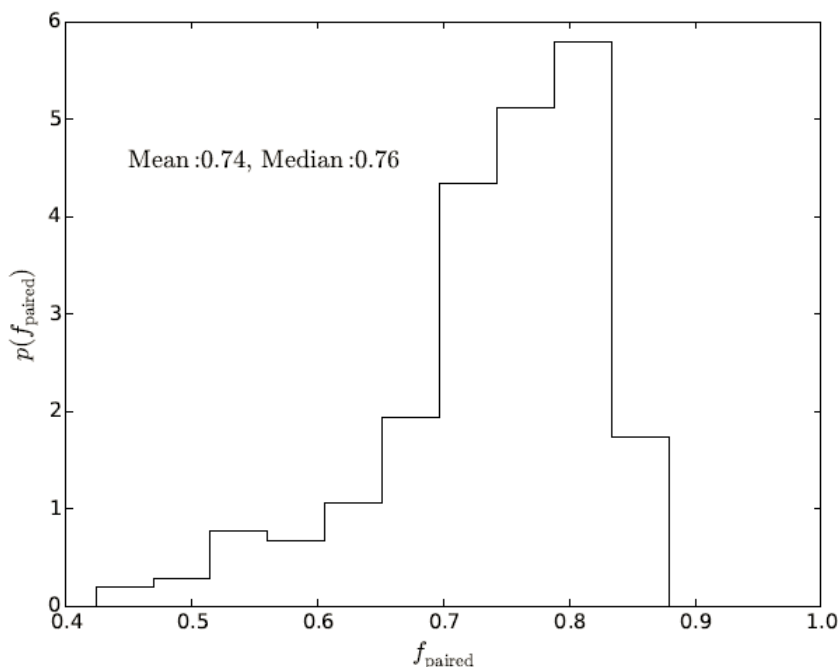
Orthog

$$x_{gas} = -0.5 * \Omega_m - \frac{C_{gas}}{\Omega_m} x_{DM}$$

$$x_{DM} = 0.5 * \Omega_b / \Omega_m x_{gas}$$

Quantifying the DM-gas segregation

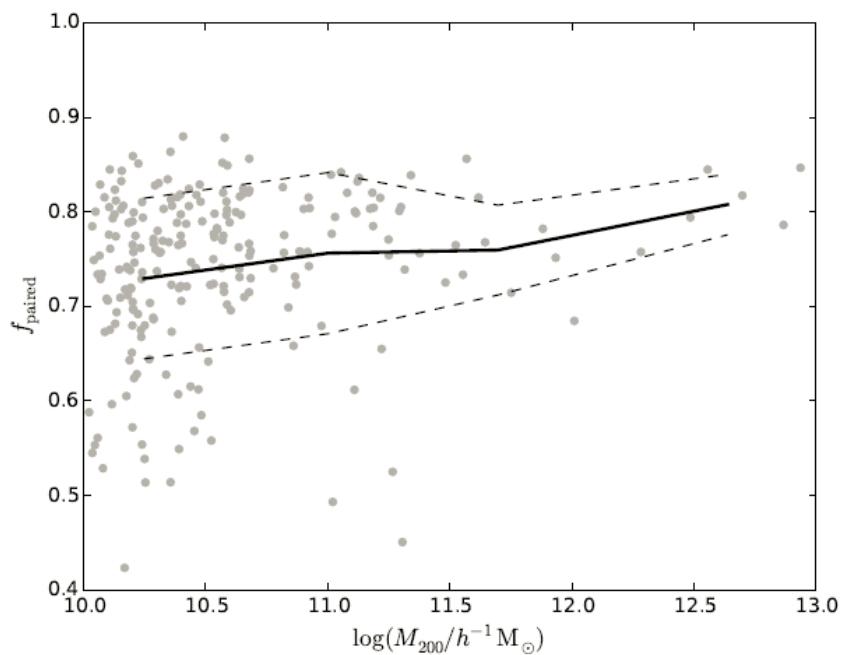
- Paired fraction $f_{\text{paired}} \equiv \frac{2N_{\text{pairs}}}{N_{\text{tot}}}$



On average 26% of the particles in a halo are initially segregated, namely from different Lagrangian region

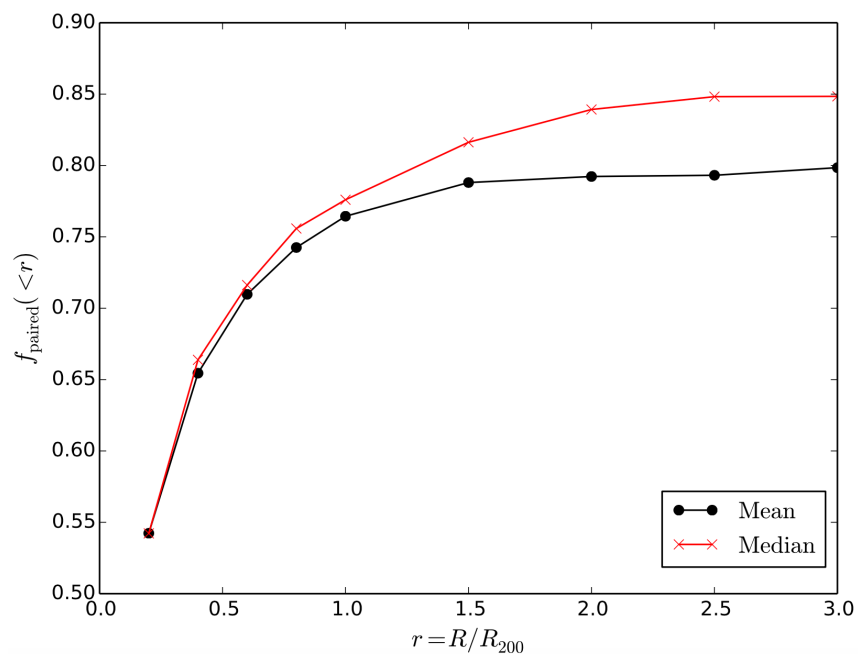
PDF of paired fraction

halo mass dependence



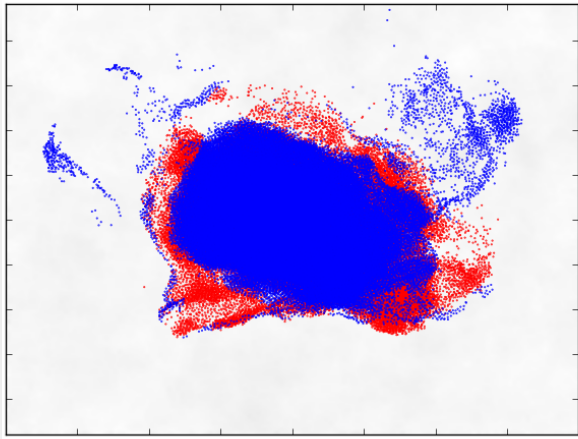
More massive halo less segregated

Radial dependence for
100 most massive halo

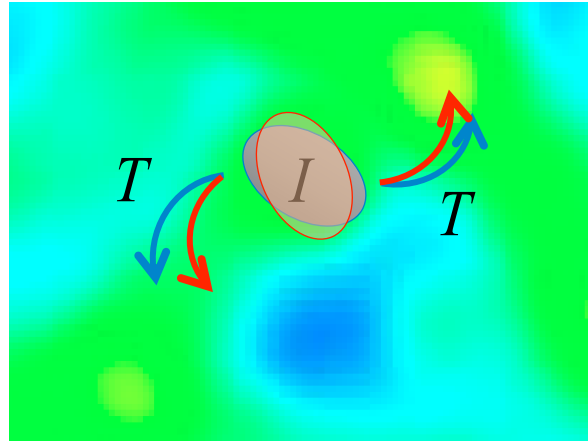


the inner particles of halo tend to be more segregated.

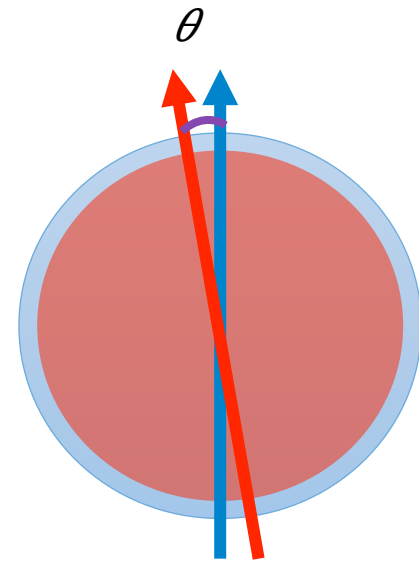
Implications on angular momentum misalignment between gas and DM



Segregation

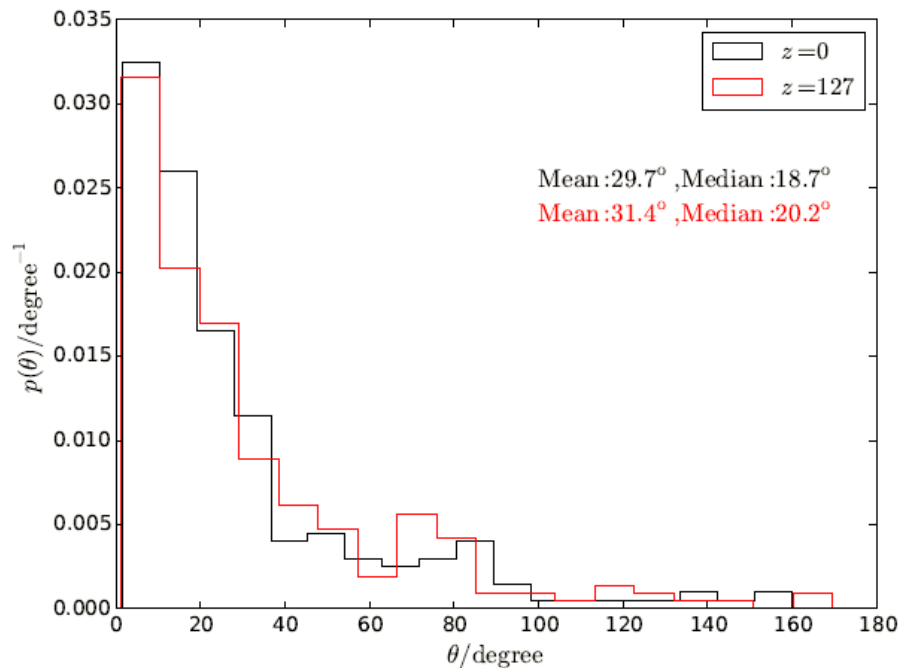


Different tidal forces

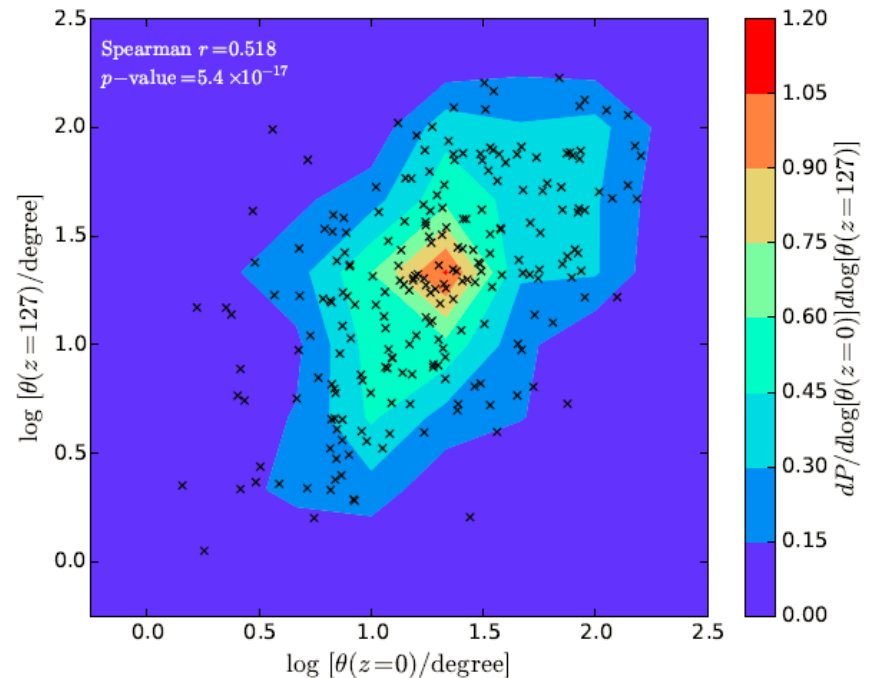


Different A.M.

Misalignment at $z=0$ and in the initial conditions



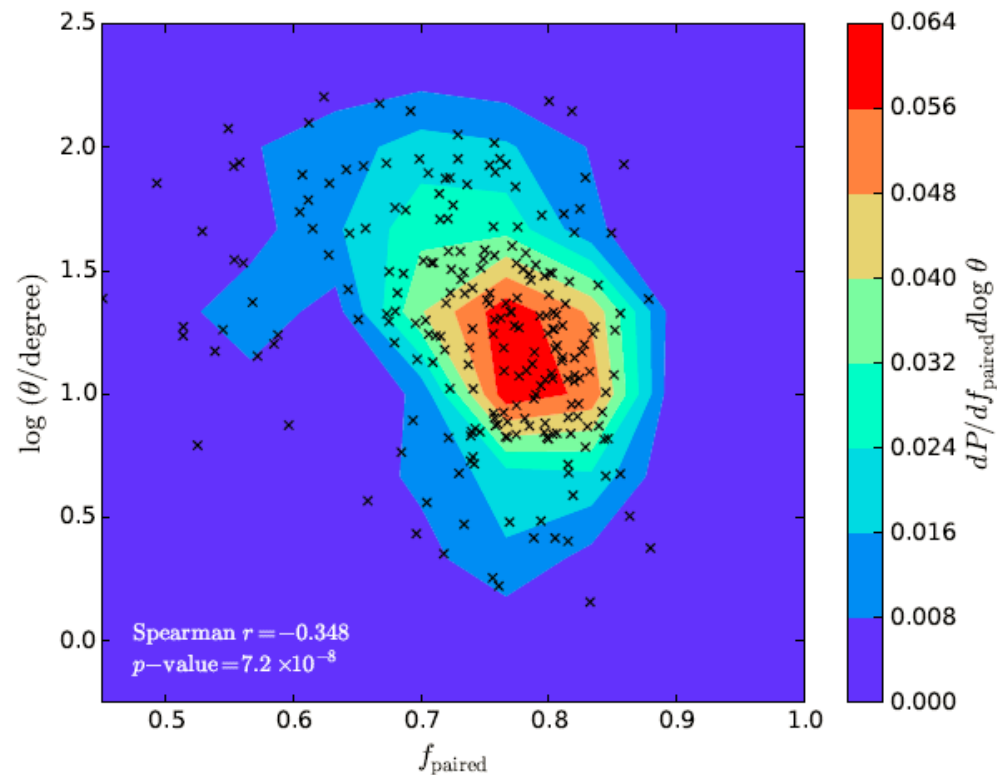
PDF misalignment angle is identical for the present day halos and their counterparts in the initial conditions



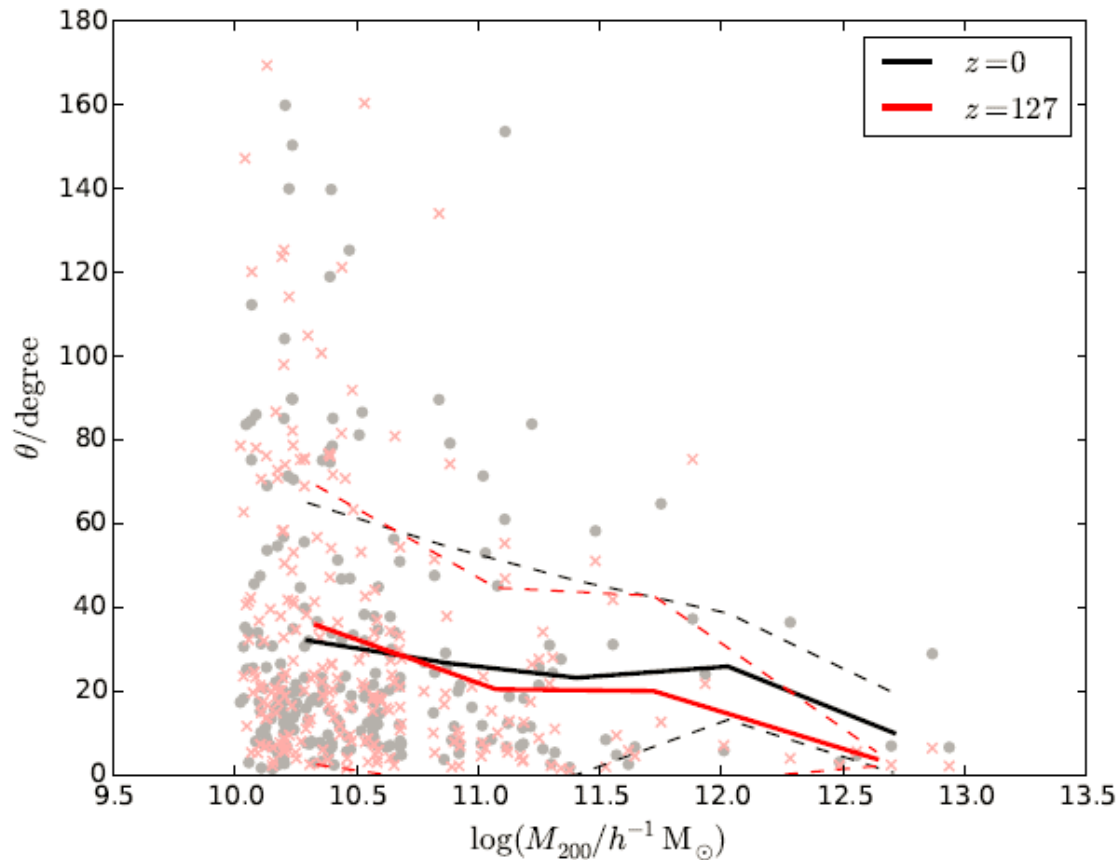
Strong correlation between Misalignment angle at $z=10$ and $z=127$ for individual halos

Correlation between the segregation strength and misalignment angle

- A halo that has a stronger segregation tends to have a larger misalignment angle



Halo mass dependence of misalignment angle naturally explained



Less massive halo, more segregated in the initial conditions

→ larger misalignment angle.

→ Misalignment between gas and dark matter should be caused by gas-dark matter segregation effect in the initial conditions.

Summary

- Dark matter and gas components of halos are initially segregated with varying extend.
- On average $\sim 25\%$ of the matter (dark and baryonic) in a halo originates from different Lagrangian region.
- The gas-DM segregation effect is at odd with the standard assumptions in semi-analytical approaches.
- The gas-DM segregation effect naturally solve the AM misalignment puzzle.